

Part. 1 Introduction				
09:00 – 09:15	Welcome and Overview of the Day			
09:15 - 09:40	Space Debris Mitigation Principles and their Effects	Holger Krag, ESA		
09:40 - 10.10	International Guidelines	Thomas Schildknecht, AIUB		
10:10 - 10:30	Current Implementation Levels	Stijn Lemmens, ESA		
10.30 - 11.00	Coffee Break			

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	Part 2 Standards and Processes	
11.00 – 11.30	The GVF Best Practices Document	Dan Oltrogge, AGI
11.30 - 11.45	Space Debris Standards on ISO Level	Dan Oltrogge, AGI
11.45 – 12.00	Debris related subordinated ISO-standards	Vitali Braun, ESA
12.00 - 13.00	Lunch Break	
13.00 - 13.30	Future Debris Management Concepts	Stijn Lemmens, ESA

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Day #2 – Part 3



Part 3 Processes and Implementation Examples				
13.30 - 14.30	The ESA Space Debris Mitigation Process, Handbooks and Examples	Rosario Nasca, ESA		
14:30 - 15:00	French Process for Debris Mitigation Compliance Verification	Laurent Francillaut, CNES		
15:00 - 15:30	The Belgium Space Debris Mitigation Process	Jean-Francois Mayence, BELSPO		
15.30 - 16.00	Coffee Break			
16:00 - 16:30	Licensing Space Activities in the era of New Space	Toby Harris, UKSA		
16:30 - 17:00	The New Zealand Process for Space Debris Mitigation	Dave Willing, New Zealand Space Agency		
17:00 - 17:30	Space Debris Mitigation – Implementation by DLR	Jan Grosser, DLR		
17:30	Adjourn	·		

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Space Debris Mitigation Principles and their Effects

Holger Krag, Head of ESA's Space Debris Office 20/03/2019

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Evolution of a Fragment Cloud - Animation





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Iridium/Cosmos Collision (Animation)





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business as usual

object count

time





Spatial Density





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HVI Impact Test





- HVI sample: impact of an
 Al-sphere of d = 1.2cm (m
 ≈ 1.7g)at v = 6.8km/s on
 an Al-block of diameter
 18cm and height 8.2cm
- Crater depths: 5.3cm

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Impact Energy









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HST Solar Array Retrieval





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Impacts in Solar Arrays

MIR Impact



Mir Station – Starboard Window Service Module Jan 1996



[Source: Thomas Reiter]

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IADC - Protected Regions



Inter-Agency Space Debris Coordination Committee



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Current Guidelines



- Drafted on request of UNCOPUOUS (presented in 2003)
 - Prevent Release of mission related objects:
 - Passivation
 - Disposal (90% reliability):
 - GEO: Graveyard orbit
 - LEO: Limit Orbital Lifetime to < 25 years after mission in LEO
 - Collision Avoidance
 - Limit Risk on-ground to 1:10.000 per re-entry event

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Effectivity of Measures

Effectiveness of Mtigation Measures





European Space Agency

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Battery Break-up Causes

Battery Failure Modes potentially leading to break-up:

- Over-temperature
- □ Short-circuit (internal or external)
- Over-charge
- Over-discharge
- Structural issues, damage



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Battery Break-up Causes





Leading to thermal runaway

- → Increase of internal pressure
- → Break-up if reaction is too quick for protections to react in time

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Electrical Passivation: Methods





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Why Propulsion Passivation?



Current technology only allows to deplete hydrazine tanks to ~5 bar and about 1% residual propellant
Main risks

1. Propellant dissociation causing tank burst

Т (°К)	T (°C)	Thermal Runaway
373	100	20 days
423	150	19 hours
473	200	1.5 hours

> 50°C, hydrazine can begin to dissociate

Reaction is exothermic

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2. Hypervelocity

explosion

impacts causing

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Disposal from LEO





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Disposal from LEO



- Up-down manoeuvre due to Earth-sensor constraints (can be designed for!)
- Estimated remaining lifetime below 15y

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European Space Agency



Drag Augmentation Devices



- Particularly attractive for satellites without a propulsion system.
- Applicable to uncontrolled reentry of satellites in orbit altitudes below 700 km.
- Stabilization of the attitude is difficult to achieve for altitudes above 550 km.



Deployed Icarus sail (on an engineering model).

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Risk on Ground





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Source: Paul Maley



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Image: Image

Critical elements in a spacecraft





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Uncontrolled vs. Controlled Re-entry





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